

AND THE FORWARD PHYSICS FACILITY

Quarks 2021, XXI International Seminar on High-Energy Physics 8 June 2021

Jonathan Feng, UC Irvine





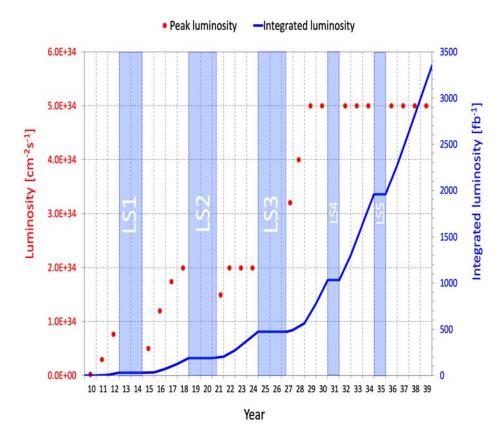






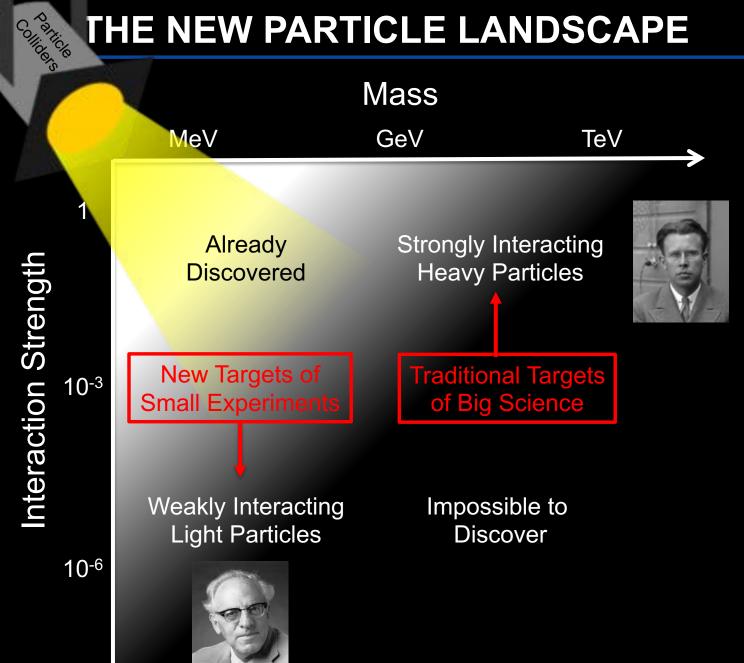
FUTURE OF PARTICLE PHYSICS

- A fact that should be better appreciated: we are in a critical period.
- In the next year or two, we will either make plans to exploit new opportunities at the HL-LHC, or lose them until 2045, 2060, 2075, ..., long after many of us are (professionally) dead.



 What should we do now? One answer: explore a rich BSM and SM physics program in the far forward region that greatly expands the LHC physics potential with relatively little additional investment.

THE NEW PARTICLE LANDSCAPE



THE THERMAL RELIC LANDSCAPE

Mass
MeV GeV TeV

Interaction Strength

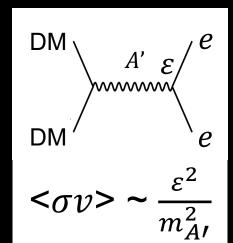
Already Discovered

Too Little to be Dark Matter

Weakly Interacting Light Particles Strongly Interacting Heavy Particles

Impossible to Discover

Too Much to be Dark Matter



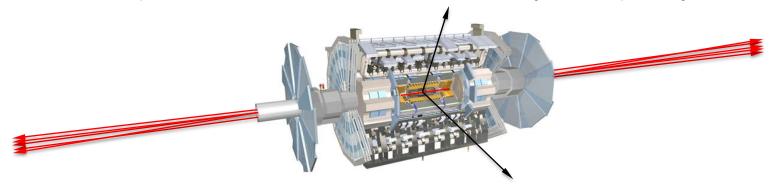
Boehm, Fayet (2003) Feng, Kumar (2008)

10-6

10-3

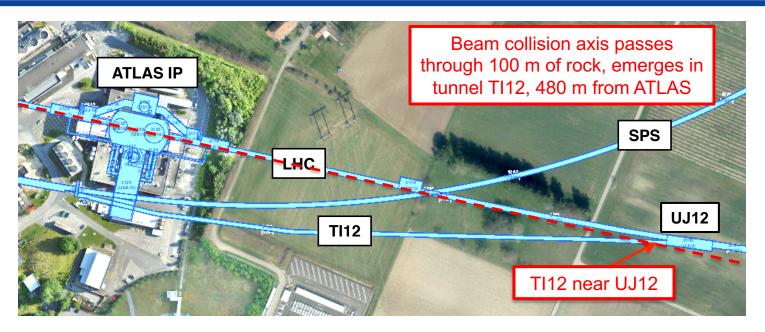
SEARCHES FOR NEW LIGHT PARTICLES

- If new particles are light and weakly interacting, existing LHC detectors are perfectly designed NOT to see them.
- Existing detectors are designed to find new heavy particles. These
 particles are produced almost at rest and decay isotropically.



- But new light particles are mainly produced in the decays of light particles: π , η , K, D and B mesons. These are mainly produced along the beamline, and so the new particles disappear through the holes that let the beams in.
- Clearly we need a detector to exploit the "wasted" $\sigma_{\text{inel}} \sim 100$ mb and cover these "blind spots" in the forward region. If we go far enough away, the proton beams are bent by magnets (it's a circular collider!), whereas the new light particles will go straight.

LOCATION, LOCATION



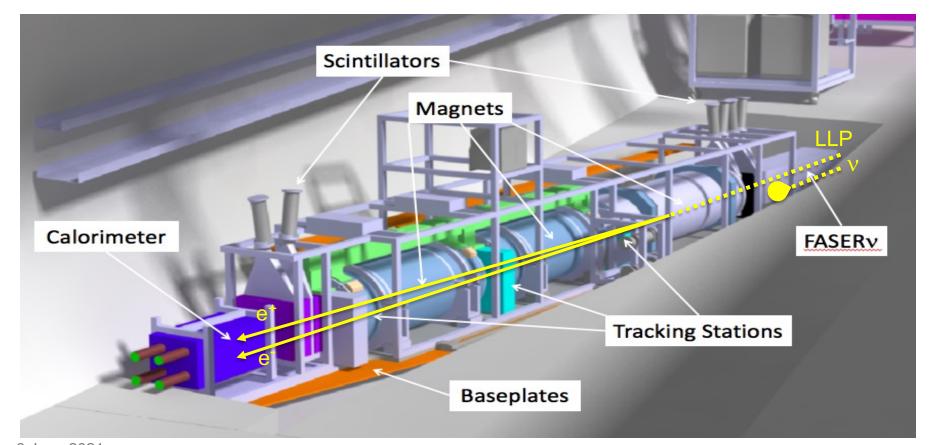


FASER TIMELINE

- September 2017: Proposed by Feng, Galon, Kling, Trojanowski.
- July 2018: Submitted LOI to CERN LHCC
- October 2018: Approval from ATLAS SCT and LHCb Collaborations for use of spare detector modules
- November 2018: Submitted Technical Proposal to LHCC
- November 2018 January 2019: Experiment funded by grants from the Heising-Simons and Simons Foundations
- March 2019: FASER fully approved by CERN along with host lab costs
- December 2019: FASERv fully approved by CERN along with host lab costs
- March 2021: FASER fully installed, commissioning of the detector begins
- April 2021: FASERv announces first collider neutrino candidates
- Early 2022: FASER and FASERv begin collecting data in Run 3

THE SIGNAL

- Nothing incoming and 2 ~TeV, opposite-sign charged tracks pointing back to the ATLAS IP: a "light shining through (100 m-thick) wall" experiment.
- Scintillators veto incoming charged tracks (muons), and permanent dipole magnets split the charged tracks, which are detected by 3 tracking stations and a calorimeter.

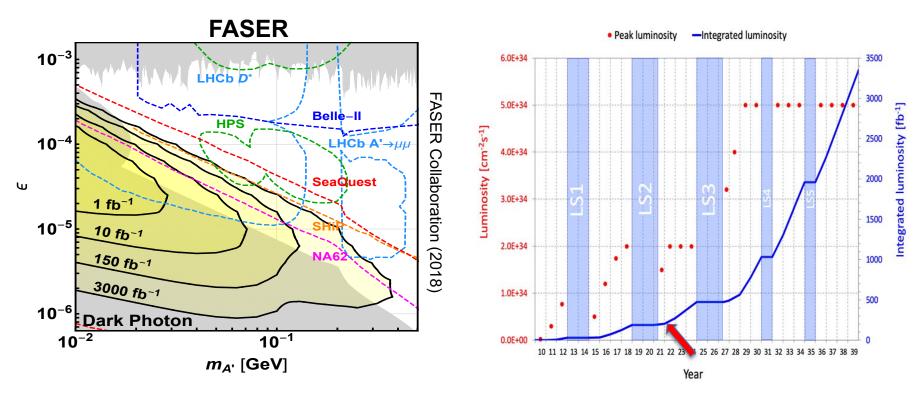


8 June 2021





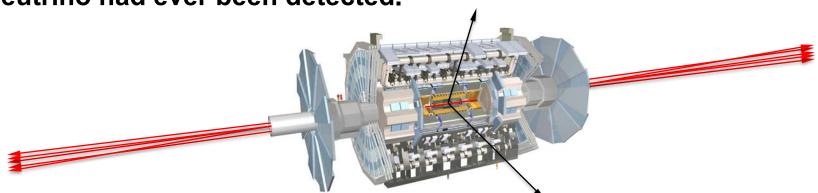
DARK PHOTON SENSITIVITY REACH



- FASER probes new parameter space with just 1 fb⁻¹ starting in 2022.
- Without detector upgrade, HL-LHC extends (L*Volume) by factor of 3000; possible upgrade to FASER 2 (R = 1m, L = 10m) extends (L*Volume) by factor of ~10⁶.

COLLIDER NEUTRINOS

- In addition to the possibility of hypothetical new light, weakly-interacting particles, there are also known light, weakly-interacting particles: neutrinos.
- But the high-energy ones, which interact most strongly, are overwhelmingly produced in the far forward direction. Before April 2021, no collider neutrino had ever been detected.



- Of course, if they can be detected, there is a fascinating new world of LHC neutrinos that can be explored.
 - The neutrino energies are ~TeV, highest human-made energies ever.
 - All flavors are produced($\pi \to \nu_{\mu}$, $K \to \nu_{e}$, $D \to \nu_{\tau}$) and both neutrinos and antineutrinos.

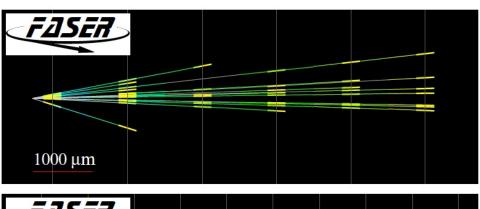
De Rujula, Ruckl (1984); Winter (1990)

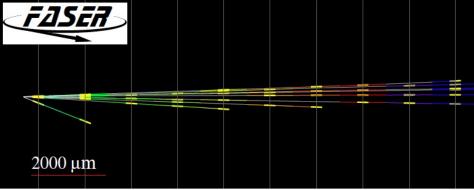
FASER Collaboration (2019); Bai, Diwan, Garzelli, Jeong, Reno (2020)

FIRST COLLIDER NEUTRINOS

- In 2018 a 29 kg FASER pilot emulsion detector collected 12.2 fb⁻¹ on the beam collision axis (installed and removed during Technical Stops).
- Based on 11kg fidiucial mass, expect ~3.3 neutrino interactions. On 14 May 2021, we announced the first direct detection of 6 collider neutrino candidates above 11 expected neutral hadron background events (2.7σ).
- Not the discovery of collider neutrinos, but a promising sign of things to come!





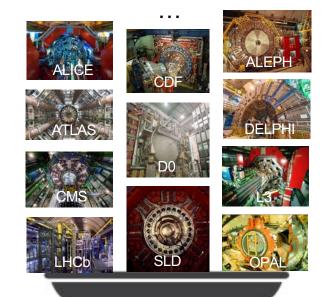


LOCATION, LOCATION

FASER Pilot Detector

Suitcase-size, 4 weeks \$0 (recycled parts)

6 neutrino candidates



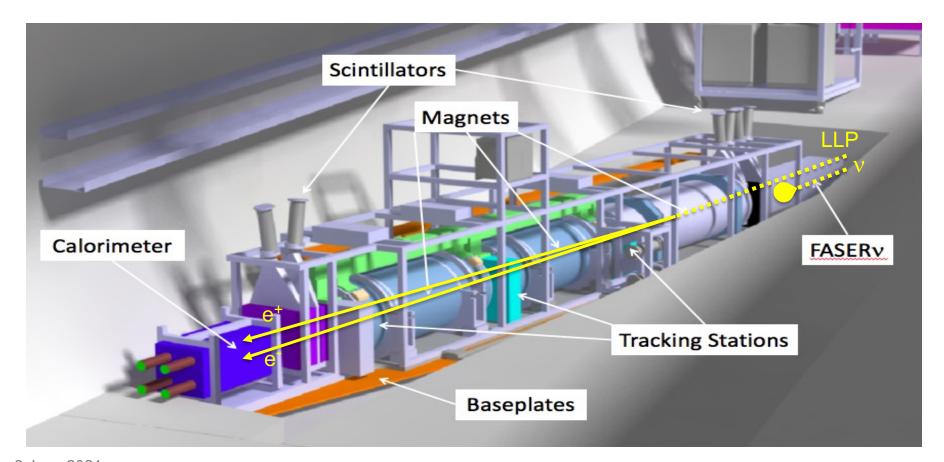
All previous collider detectors

Building-size, decades ~\$109

0 neutrino candidates

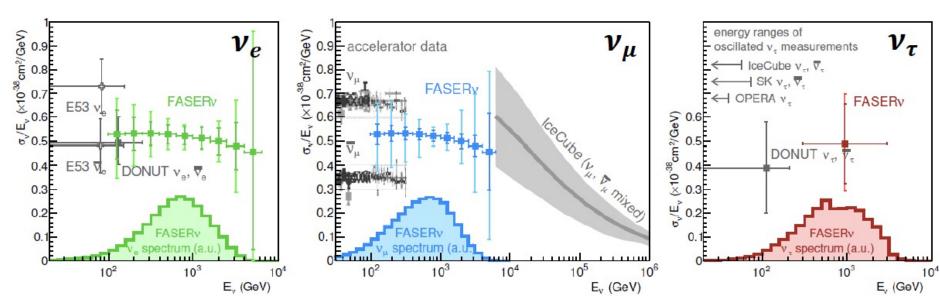
THE FASERV DETECTOR

- FASERv is designed to detect neutrinos of all flavors.
 - 25cm x 30cm x 1.1m detector consisting of 770 emulsion layers interleaved with 1mm-thick tungsten plates; target mass = 1.1 tonne.
 - Emulsion swapped out every ~10-30 fb⁻¹, total 10 sets of emulsion for Run 3.



NEUTRINO PHYSICS

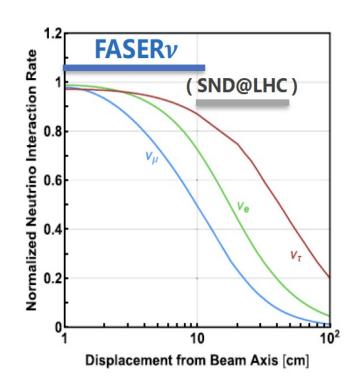
- In Run 3 (2022-24), FASERv will
 - Detect the first collider neutrino.
 - Record ~1000 ν_e , ~10,000 ν_μ , and ~10 ν_τ interactions at TeV energies, the first direct exploration of this energy range for all 3 flavors.
 - Double the world's supply of tau neutrinos.
 - Distinguish muon neutrinos from anti-neutrinos by combining FASER and FASERv data, and so measure their cross sections independently.



FASER Collaboration 1908.02310 (2019)

QCD PHYSICS

- The forward production of hadrons is currently subject to large uncertainties. Forward v experiments will provide useful insights.
 - On- and off-axis neutrino detectors provide complementary information $(\pi \to \nu_{\mu}, K \to \nu_{e}, D \to \nu_{\tau}).$
 - Different target nuclei (lead, tungsten) probe different nuclear pdfs.
 - Strange quark pdf through $vs \rightarrow lc$.
 - Forward charm production, intrinsic charm.
 - Refine simulations that currently vary greatly (EPOS-LHC, QGSJET, DPMJET, SIBYLL, PYTHIA...).
 - Provide essential input to astroparticle experiments; e.g., distinguish galactic neutrino signal from atmospheric neutrino background at IceCube.

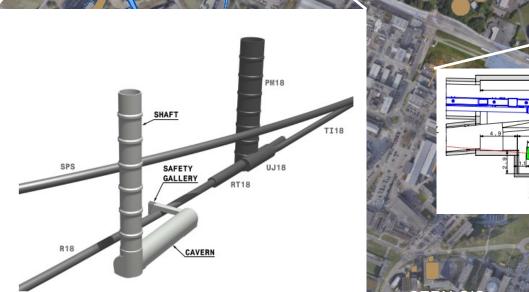


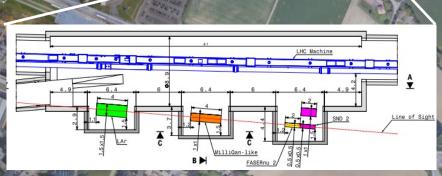
FORWARD PHYSICS FACILITY

- FASER, FASERv, and other proposed far-forward detectors are currently highly constrained by 1980's infrastructure that was never intended to support experiments.
- At the same time, it is becoming clear that there is a rich physics program in the far-forward region.
 - New particle searches, neutrinos, QCD, MC event generators, milli-charged particles, dark matter, dark sector, cosmic neutrinos, ...
- Strongly motivates creating a dedicated facility to house far-forward experiments for the HL-LHC era from 2027-37.
 - Snowmass LOI: 240 authors from many different communities
 - FPF Kickoff Meeting: 9-10 Nov 2020, https://indico.cern.ch/event/955956,
 2nd FPF Workshop: 27-28 May 2021, https://indico.cern.ch/event/1022352.
 - 1st paper on the physics potential of the FPF will be completed July 2021,
 Snowmass white paper to be completed February 2022.

FPF LOCATION

Possibilities under active investigation: enlarge existing cavern UJ12, 480 m from ATLAS and shielded from the ATLAS IP by ~100 m of rock; or create a new shaft and cavern ~612 m from ATLAS past UJ18.





See Kincso Balazs talk, FPF2 Meeting

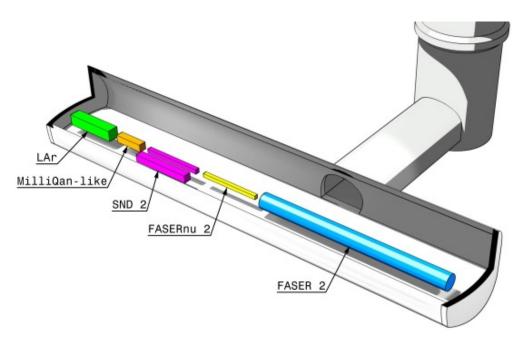
SPS

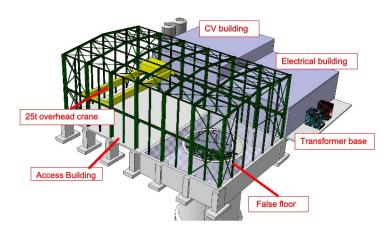
CERN GIS

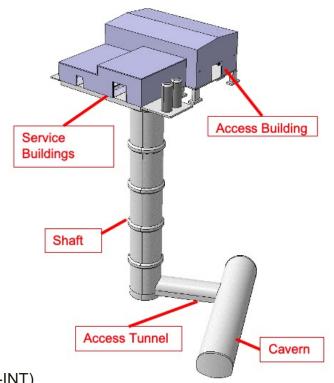
FPF: NEW SHAFT AND CAVERN

Many advantages

- Construction access far easier
- Access possible during LHC operations
- Designed around needs of experiments
- Very preliminary (Class 4) cost estimate:
 ~\$23 MCHF.







Design by Liam Dougherty (EN-ACE-INT)

NEW PHYSICS SEARCHES AT THE FPF

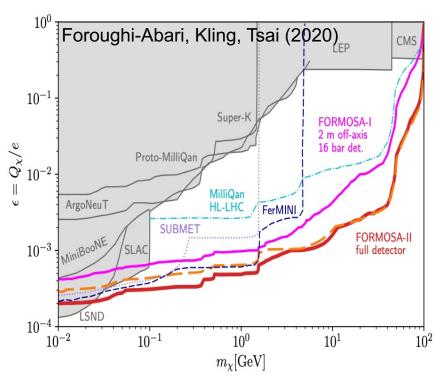
- The FPF will house a number of experiments
- FASER 2, an upgraded
 FASER with R = 1 m, L = 10
 m, can discover all
 candidates with
 renormalizable couplings
 (dark photon, dark Higgs,
 HNL); ALPs with all types of
 couplings (γ, f, g); and many
 other particles.
- Other experiments can probe neutrinos and many other interesting ideas.

Benchmark Model	Underway	FPF
BC1: Dark Photon	FASER	FASER 2
BC1': U(1) _{B-L} Gauge Boson	FASER	FASER 2
BC2: Dark Matter	-	FLArE
BC3: Milli-Charged Particle	-	FORMOSA
BC4: Dark Higgs Boson	-	FASER 2
BC5: Dark Higgs with hSS	-	FASER 2
BC6: HNL with e	-	FASER 2
BC7: HNL with μ	-	FASER 2
BC8: HNL with τ	FASER	FASER 2
BC9: ALP with photon	FASER	FASER 2
BC10: ALP with fermion	FASER	FASER 2
BC11: ALP with gluon	FASER	FASER 2

BC2: MILLI-CHARGED PARTICLES

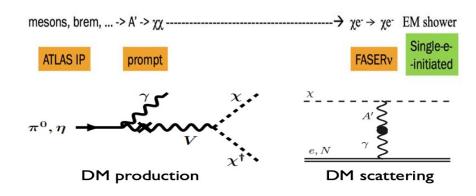
- Currently the target of the MilliQan experiment near the CMS IP.
- MilliQan Demonstrator (Proto-MilliQan) already probes new region. Full
 MilliQan planned to run in this location at HL-LHC, but the sensitivity can
 be improved significantly by moving it to the FPF (FORMOSA).



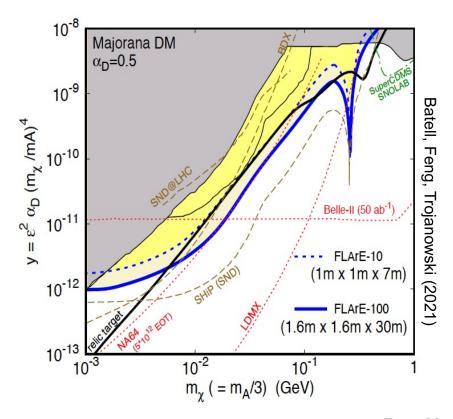


BC3: DARK MATTER

- If m_{LLP} > 2m_{DM}, the LLP will typically decay to dark matter, leading to a highly collimated beam of dark matter particles.
- Can look for the resulting DM to scatter off electrons at FLArE, Forward Liquid Argon Experiment, a proposed 10 to 100 tonne LArTPC.



FLArE probes most of the favored/ allowed relic target region. Complementary to missing energy experiments that probe more of the "too large $\Omega_{\chi}h^2$ " region, but don't detect DM scattering.



SUMMARY

- New target for discovery: neutrinos and light and weakly-interacting particles probed by fast, small, cheap experiments in the far-forward region of the LHC.
- FASER and FASERv: 3.5 years from theory proposal to completion, 5 m long, ~\$2M. Data-taking starts in 2022 with a rich physics program.
 - BSM: searches for dark photons, HNLs, ALPs, new forces.
 - SM: opens the new field of LHC neutrino physics. ~1000 ν_e , ~10,000 ν_μ , ~10 ν_τ at TeV energies. Implications for neutrino properties, forward hadron production, cosmic ray and cosmic neutrino physics.
- Forward Physics Facility: Proposed facility to house a suite of farforward experiments for the HL-LHC era from 2027-37.
- FLArE: Forward Liquid Argon Experiment, proposed 10-100 tonne LArTPC for dark matter searches, neutrino physics.